

# Resonant Leptogenesis in B-L Model with modular symmetry

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- Flavor Symmetry

# Resonant Leptogenesis

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# Baryon Asymmetry

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There are baryon asymmetry in universe.

$$5.8 \cdot 10^{-10} < \frac{n_b - n_{\bar{b}}}{n_\gamma} < 6.5 \cdot 10^{-10}$$

The origin of baryon asymmetry is mystery.

# Sakharov's condition

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The baryon asymmetry can be dynamically generated ('baryogenesis') provided that

1. B violation
2. C and CP violation
3. Out of thermal equilibrium

# Baryogenesis in SM

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## Electroweak baryogenesis

- B violation
  - sphaleron process
- C and CP violation
  - CKM matrix
- Out of thermal equilibrium
  - 1<sup>st</sup> order phase transition

Can electroweak baryogenesis in SM produce a sufficiently large asymmetry?

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NO!

# Can electroweak baryogenesis in SM produce a sufficiently large asymmetry?



NO!

The SM fails on two aspects:

1. The Higgs sector does not give a strongly first order PT.

Requirement:  $m_H < 70\text{GeV}$       LHC:  $m_H = 125\text{ GeV}$

2. CP Violation in CKM is not enough.

Requirement:  $Y_B \sim 10^{-10}$       CKM:  $\frac{n_B - n_{\bar{B}}}{n_B + n_{\bar{B}}} \sim 10^{-20}$

# Baryogenesis Via Leptogenesis

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SM + right-handed Majorana neutrinos

- B violation

- L violation by majorana neutrinos decay

- Sphaleron process

- C and CP violation

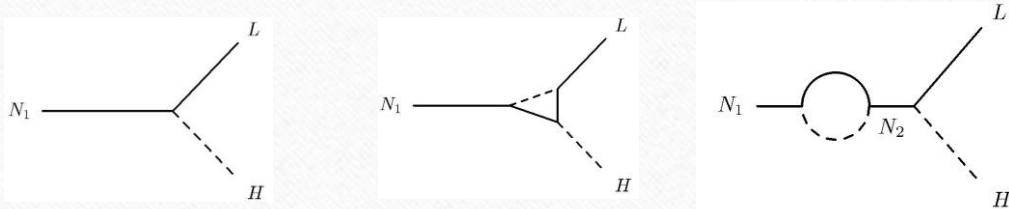
- Neutrino mixing matrix (complex)

- Out of thermal equilibrium

- Out of equilibrium decay of right-handed neutrino

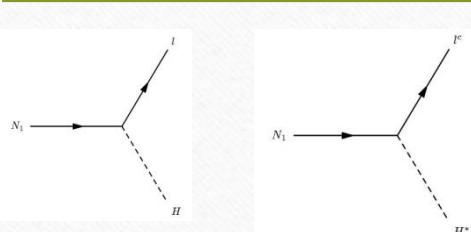
# Baryogenesis Via Leptogenesis

- L asymmetry is generated due to CP asymmetry that arises through interference of tree level and one-loop diagrams.



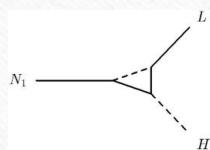
- Sphaleron process converts L asymmetry into B asymmetry.

# CP Asymmetry



Right-handed Majorana neutrino can decay into lepton and anti-lepton.

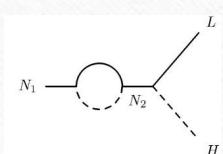
## Vertex contribution



$$V_j = 2 \frac{M_j^2}{M_1^2} \left[ \left( 1 + \frac{M_j^2}{M_1^2} \right) \log \left( 1 + \frac{M_1^2}{M_j^2} \right) - 1 \right]$$

$$\begin{aligned} \epsilon_1 &\equiv \frac{\sum_j [\Gamma(N_1 \rightarrow \ell_j h) - \Gamma(N_1 \rightarrow \ell_j^C h^*)]}{\sum_j [\Gamma(N_1 \rightarrow \ell_j h) + \Gamma(N_1 \rightarrow \ell_j^C h^*)]} \\ &= - \sum_{j=2,3} \frac{M_1}{M_j} \frac{\Gamma_j}{M_j} \left( \frac{V_j}{2} + S_j \right) \frac{\text{Im} [(Y_D^2)_{1j}^2]}{|Y_D^2|_{11} |Y_D^2|_{jj}} \quad \frac{\Gamma_j}{M_j} = \frac{|Y_D^2|_{jj}}{8\pi} \end{aligned}$$

## Self-energy contribution



$$S_j = \frac{M_j^2 \Delta M_{1j}^2}{(\Delta M_{1j}^2)^2 + M_1^2 \Gamma_j^2}$$

$$\Delta M_{1j}^2 = M_j^2 - M_1^2$$

# Hierarchical case

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If right-handed neutrinos have a hierarchical mass spectrum ( $M_{2,3} \gg M_1$ ), we can write a CP asymmetry parameter as

$$\epsilon_1 \sim \frac{3}{16\pi} \frac{m_\nu M_1}{v^2} \sin \delta \sim 10^{-6} \frac{m_\nu}{0.05\text{eV}} \frac{M_1}{10^{10}\text{GeV}} \sin \delta$$

The present Baryon asymmetry

$$Y_B \sim \kappa \frac{\epsilon_1}{g_*} \sim 10^{-10}$$

$g_* = \mathcal{O}(100)$  is the number of relativistic degrees of freedom.

In the SM  $\kappa \sim 2 \times 10^{-2} \left( \frac{0.05\text{eV}}{m_\nu} \right)^{1.1}$

# Majorana mass bound

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$$Y_B \sim 10^{-10} \left( \frac{0.05\text{eV}}{m_\nu} \right)^{0.1} \frac{M_1}{10^{10}\text{GeV}} \sin \delta \sim 10^{-10}$$

$$m_\nu \sim \sqrt{m_{atm}^2} \sim 0.05\text{eV}$$

The lightest Majorana mass is heavier than  $10^{10}\text{GeV}$  .

# Resonant Leptogenesis

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- The Majorana masses are heavier than  $10^{10} GeV$ , if the spectrum of Majorana masses has hierarchy.
- If the Majorana mass of right-handed neutrino is smaller than a few TeV, general leptogenesis can not work.

# Resonant Leptogenesis

- The Majorana masses are heavier than  $10^{10} GeV$ , if the spectrum of Majorana masses has hierarchy.
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Resonant-Leptogenesis

# Resonant Leptogenesis

When two right-handed neutrinos have mass differences comparable to their decay widths ( $M_2^2 - M_1^2 \sim M_1 \Gamma_2$ ), self-energy correction dominates.

$$\begin{aligned}\epsilon_1 &\sim \frac{\text{Im}[Y^2]_{12}^2}{|Y^2|_{11}|Y^2|_{22}} \frac{M_1 \Gamma_2 (M_2^2 - M_1^2)}{(M_2^2 - M_1^2)^2 + M_1^2 \Gamma_2^2} \\ &\sim \frac{1}{2} \frac{\text{Im}[Y^2]_{12}^2}{|Y^2|_{11}|Y^2|_{22}}\end{aligned}$$

$\epsilon_1$  can be even O(1).

$$M_1 \sim TeV, \quad \frac{M_2^2 - M_1^2}{M_1^2} \sim 10^{-12}$$

# Degenerate masses

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$M_1 = M_2$  @ High scale

RG evolution

$$\begin{aligned} \longrightarrow M_2^2 - M_1^2 &\propto Y_\nu^\dagger Y_\nu (y v_{B-L})^2 \\ &\propto \Gamma M \end{aligned}$$

Flavor symmetry

$S_4$  Modular symmetry

# Flavor Symmetry

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# Model

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3 Right-Handed Neutrinos  $\Rightarrow$  Resonant Leptogenesis

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Dark Matter candidate

$U(1)$  B-L Gauge Symmetry  $\Rightarrow$  Thermal Production of Neutrinos

$S_4$  Modular Symmetry  $\Rightarrow$  Resonant Leptogenesis

# Modular Symmetry

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$$\tau \rightarrow \gamma\tau = \frac{a\tau + b}{c\tau + d}, \text{ where } a, b, c, d \in \mathbb{Z} \text{ and } ad - bc = 1, \text{ Im}[\tau] > 0$$

Generators

$$S : \tau \rightarrow -\frac{1}{\tau},$$

$$T : \tau \rightarrow \tau + 1$$

$$S^2 = I, (ST)^3 = I$$

$$\Gamma(N) = \left\{ \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in SL(2, \mathbb{Z}) , \quad \begin{pmatrix} a & b \\ c & d \end{pmatrix} \equiv \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \pmod{N} \right\}$$

$$T^N = I$$

$$\Gamma(4) \approx S_4$$

# $S_4$ symmetry

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All permutations among four objects

$$(x_1, x_2, x_3, x_4) \rightarrow (x_i, x_j, x_k, x_l)$$

S:

$$(x_1, x_2, x_3, x_4) \rightarrow (x_2, x_1, x_3, x_4)$$

T:

$$(x_1, x_2, x_3, x_4) \rightarrow (x_2, x_3, x_4, x_1)$$

$$S^2 = T^4 = (ST)^3 = 1$$

Representations : 1, 1', 2, 3, 3'

# $S_4$ modular symmetry

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$$S^2 = T^4 = (ST)^3 = 1 \quad \Gamma_4 \approx S_4$$

Representations : 1, 1', 2, 3, 3'

3(3') → 3 generations

2 → 2 degenerate masses

$\Gamma_2 \approx S_3$ : 1, 1', 2

$\Gamma_3 \approx A_4$ : 1, 1', 1'', 3

Minimum symmetry is  $S_4$

# Modular Symmetry

Modular form

$$f(\gamma\tau) = (c\tau + d)^k f(\tau), \quad \gamma \in \Gamma(N) \quad k : \text{Modular weight}$$

$$\begin{aligned} a = d = -1, b = c = 0 \\ f(\tau) = (-1)^k f(\tau) \end{aligned}$$

Odd modular weight  $\Rightarrow$  Vanish

Field

Modular Weight	Representation
2	2, 3'
4	1, 2, 3, 3'
6	1, 1', 2, 3, 3'

$$\phi^{(I)} \rightarrow (c\tau + d)^{-k_I} \rho^{(I)}(\gamma) \phi^{(I)}, \quad \rho^{(I)}(\gamma) : \text{Unitary representation matrix of } S_4$$

# Model

Yukawa

$$\mathcal{L}_Y = \alpha \bar{L}_L Y_{\mathbf{3}'}^{(2)} e_{R_e} H_d + \beta \bar{L}_L Y_{\mathbf{3}'}^{(2)} \ell_R H_d + \bar{L}_L \left( \gamma_1 Y_{\mathbf{3}}^{(4)} + \gamma_2 Y_{\mathbf{3}'}^{(4)} \right) N_R H_u + \delta \bar{N}_{R_1}^c Y_{\mathbf{1}}^{(6)} N_{R_1} \phi + y \bar{N}_R^c N_R \phi$$

Vanishing terms

$$\bar{L}_L N_{R_1} H_u$$



Yukawa coupling with odd modular weight

$$\bar{N}_R^c N_{R_1} \phi$$

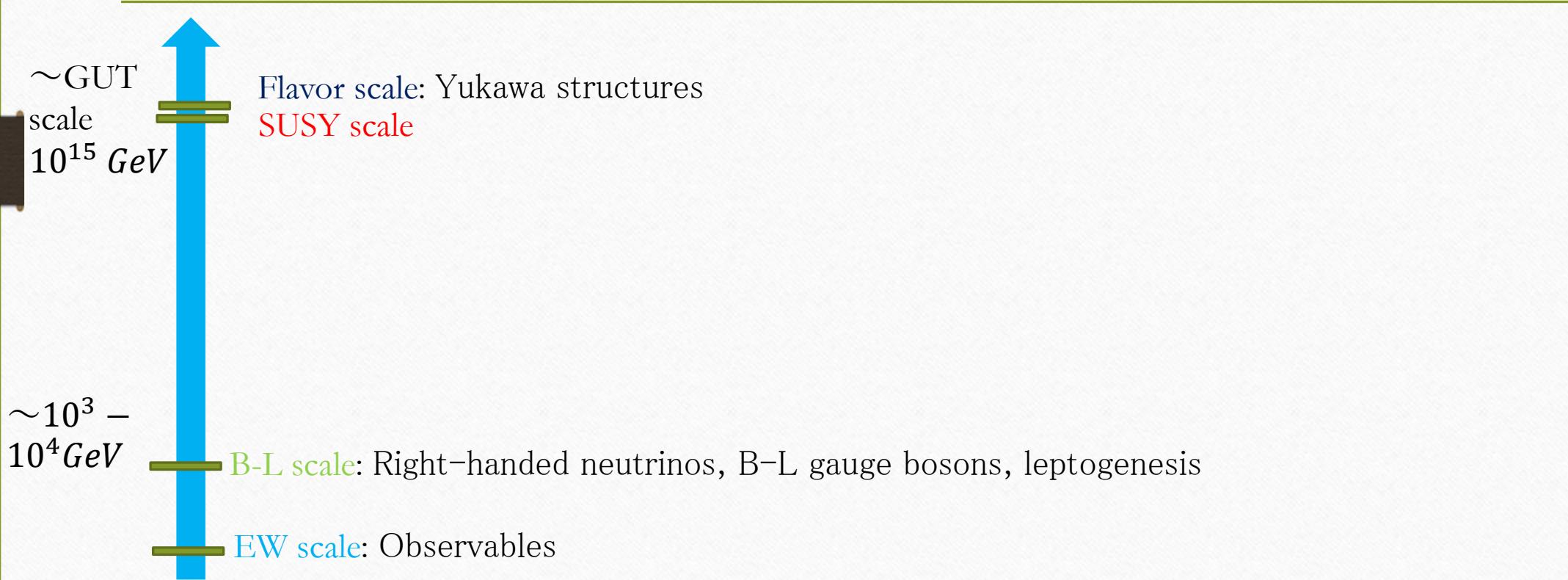
$N_{R_1}$  is stable



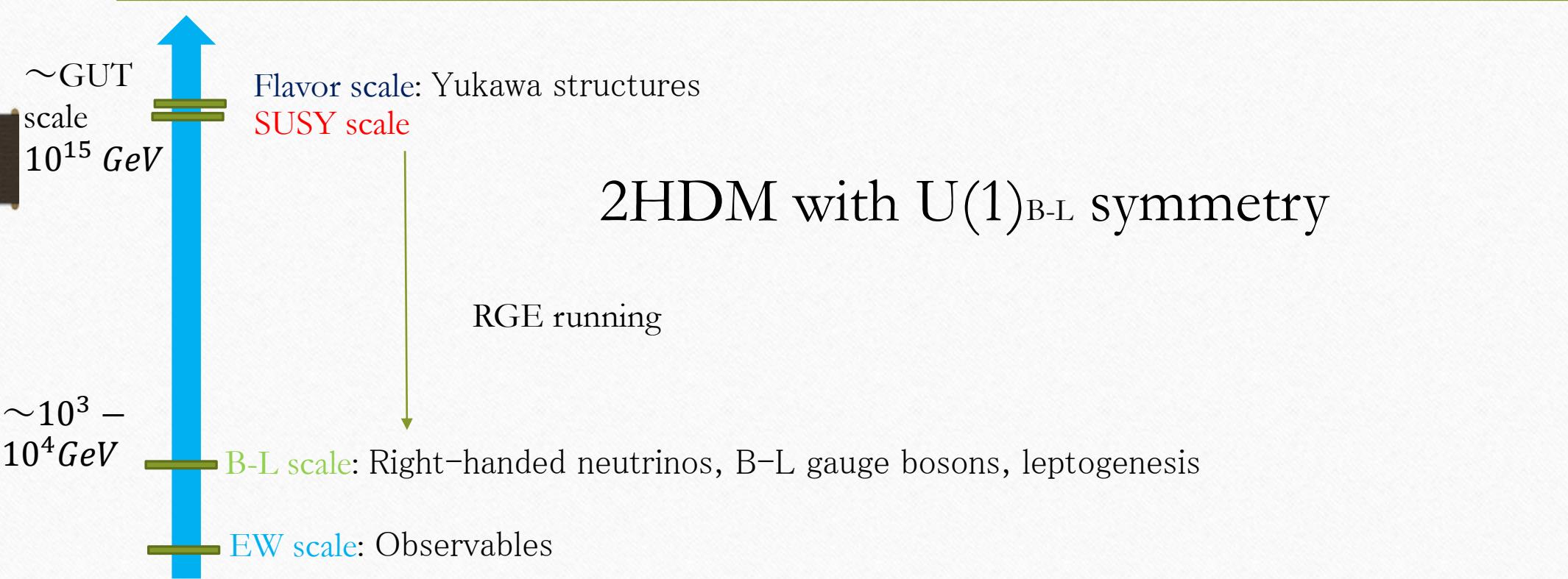
Dark Matter Candidate

	Fermions					Bosons		
	$\bar{L}_L \equiv (\bar{L}_{L_e}, \bar{L}_{L_\mu}, \bar{L}_{L_\tau})^T$	$e_{R_e}$	$\ell_R \equiv (e_{R_\mu}, e_{R_\tau})^T$	$N_{R_1}$	$N_R \equiv (N_{R_2}, N_{R_3})^T$	$H_u$	$H_d$	$\phi$
$SU(2)_L$	<b>2</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>2</b>	<b>2</b>	<b>1</b>
$U(1)_Y$	$\frac{1}{2}$	-1	-1	0	0	$-\frac{1}{2}$	$\frac{1}{2}$	0
$U(1)_{B-L}$	1	-1	-1	-1	-1	0	0	2
$S_4$	<b>3'</b>	<b>1</b>	<b>2</b>	<b>1</b>	<b>2</b>	<b>1</b>	<b>1</b>	<b>1</b>
$-k$	-1	-1	-1	-3	0	-3	0	0

# Supersymmetric B-L extended SM with $S_4$ Modular symmetry



# Supersymmetric B-L extended SM with $S_4$ Modular symmetry



# Model

	Fermions						Bosons		
	$\bar{L}_L \equiv (\bar{L}_{L_e}, \bar{L}_{L_\mu}, \bar{L}_{L_\tau})^T$	$e_{R_e}$	$\ell_R \equiv (e_{R_\mu}, e_{R_\tau})^T$	$N_{R_1}$	$N_R \equiv (N_{R_2}, N_{R_3})^T$	$H_u$	$H_d$	$\phi$	
$SU(2)_L$	<b>2</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>2</b>	<b>2</b>	<b>1</b>	
$U(1)_Y$	$\frac{1}{2}$	-1	-1	0	0	$-\frac{1}{2}$	$\frac{1}{2}$	0	
$U(1)_{B-L}$	1	-1	-1	-1	-1	0	0	2	
$S_4$	<b>3'</b>	<b>1</b>	<b>2</b>	<b>1</b>	<b>2</b>	<b>1</b>	<b>1</b>	<b>1</b>	
$-k$	-1	-1	-1	-3	0	-3	0	0	

Yukawa

$$\mathcal{L}_Y = \alpha \bar{L}_L Y_{\mathbf{3}'}^{(2)} e_{R_e} H_d + \beta \bar{L}_L Y_{\mathbf{3}'}^{(2)} \ell_R H_d + \bar{L}_L \left( \gamma_1 Y_{\mathbf{3}}^{(4)} + \gamma_2 Y_{\mathbf{3}'}^{(4)} \right) N_R H_u + \delta \bar{N}_{R_1}^c Y_{\mathbf{1}}^{(6)} N_{R_1} \phi + y \bar{N}_R^c N_R \phi$$

# Model

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Boundary conditions

$$Y_M(m_0) = \begin{pmatrix} y & 0 \\ 0 & y \end{pmatrix} \quad m_0: \text{Breaking scale of the flavor symmetry}$$

$$Y_\nu(m_0) = \begin{pmatrix} \gamma_2 y'_1 & -\gamma_1 y_1 \\ \frac{\sqrt{3}}{2} \gamma_1 y_2 - \frac{1}{2} \gamma_2 y'_3 & \frac{1}{2} \gamma_1 y_3 + \frac{\sqrt{3}}{2} \gamma_2 y'_2 \\ \frac{\sqrt{3}}{2} \gamma_1 y_3 - \frac{1}{2} \gamma_2 y'_2 & \frac{1}{2} \gamma_1 y_2 + \frac{\sqrt{3}}{2} \gamma_2 y'_3 \end{pmatrix}$$

$$Y_{\mathbf{3}}^{(4)} = \begin{pmatrix} y_1 \\ y_2 \\ y_3 \end{pmatrix}, \quad Y_{\mathbf{3}'}^{(4)} = \begin{pmatrix} y'_1 \\ y'_2 \\ y'_3 \end{pmatrix}$$

# Renormalization group equations

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$$16\pi^2 \mu \frac{dY_\nu^\dagger Y_\nu}{d\mu} = \frac{1}{2} Y_M Y_M^* Y_\nu^\dagger Y_\nu + \frac{1}{2} Y_\nu^\dagger Y_\nu Y_M Y_M^* + Y_\nu^\dagger Y_\nu (6y_t^2 - \frac{3}{2}g_1^2 - \frac{9}{2}g_2^2 - 12g_{B-L}^2 - \frac{3}{2}g_{mix}^2 - 6g_{B-L}g_{mix})$$

$$16\pi^2 \mu \frac{dY_M}{d\mu} = \underline{Y_M Y_\nu^\dagger Y_\nu + Y_\nu^\dagger Y_\nu Y_M} + Y_M Y_M^* Y_M + \frac{1}{2} Y_M Tr [Y_M Y_M^\dagger] - 6g_{B-L}^2 Y_M$$

Mass difference

Off diagonal

# Result

## parameters

Guage

$$v_{B-L} = 17 \text{ TeV}$$

$$g_{B-L} = 0.20$$

Modular

$$\tau = 0.25 + 2.5 i$$

Yukawa

$$y = 0.5$$

## predictions

Baryon number

$$\frac{n_b - n_{\bar{b}}}{n_\gamma} = 5.93 \cdot 10^{-10}$$

Active neutrino

$$m_1 = 8.46 \cdot 10^{-3} \text{ eV}$$

$$m_2 = 4.93 \cdot 10^{-2} \text{ eV}$$

Majorana neutrino

$$M_1 = 7.79 \text{ TeV}$$

$$\frac{M_2^2 - M_1^2}{M_1^2} = 9.92 \cdot 10^{-12}$$

Gauge boson

$$M_{Z'} = 6.90 \text{ TeV}$$

## Observed values

$$5.8 \cdot 10^{-10} < \frac{n_b - n_{\bar{b}}}{n_\gamma} < 6.5 \cdot 10^{-10} \text{ (95 \% CL)}$$

$$8.26 \cdot 10^{-3} \text{ eV} < m_1 < 8.97 \cdot 10^{-3} \text{ eV (3 } \sigma)$$

$$4.93 \cdot 10^{-2} \text{ eV} < m_2 < 5.10 \cdot 10^{-2} \text{ eV (3 } \sigma)$$

# Conclusions

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- Resonant leptogenesis scenario in TeV scale B-L model
- $S_4$  Modular symmetry naturally realizes degenerate masses
- RG evolutions generate a splitting of Majorana masses
- CP asymmetry parameter is automatically enhanced by resonance
- Right amount of baryon number can be obtained

# Future Work

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- Search wide region
- Quark and Charged lepton sector