

Constructing Symmetric Teleparallel theories with the Gauss-Bonnet invariant

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based on:

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arXiv: 2308.07299 - PRD (in press)



Funded by
the European Union



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CAS-JSPS-IBS CTPU-CGA Autumn Workshop on Cosmology, Gravity, Particle
Physics

Prague, October 10th, 2023

Outline

- Motivations
- Introduction
 - Theories with the Gauss-Bonnet invariant
 - Teleparallel gravity
- Gauss-Bonnet invariant in General Teleparallel formulation
 - Bulk and boundary splits
 - Flat FLRW
- Example: scalar-Symmetric-Teleparallel-Gauss-Bonnet theory

Why modified gravity?

- Account for the “invisible” ingredients: Dark Energy, Dark Matter
- H_0 and σ_8 tensions
- Singularities
- Quantum gravity
- Strong gravity regime needs to be tested
- A good way to understand GR is to modify it

General Relativity - Assumptions

General Relativity is based upon different assumptions that can be understood as the fulfilling of the Lovelock's theorem. Some assumptions are:

- **Equivalence principle**
- **General covariance:** Invariant under diffeomorphisms and Local Lorentz transformations.
- **Riemannian geometry:** The connection is the Levi-Civita one.
- **4-dimensions**
- **2nd order derivatives:** gravitational action contains only second derivatives.
- **Locality**

How to modify it?

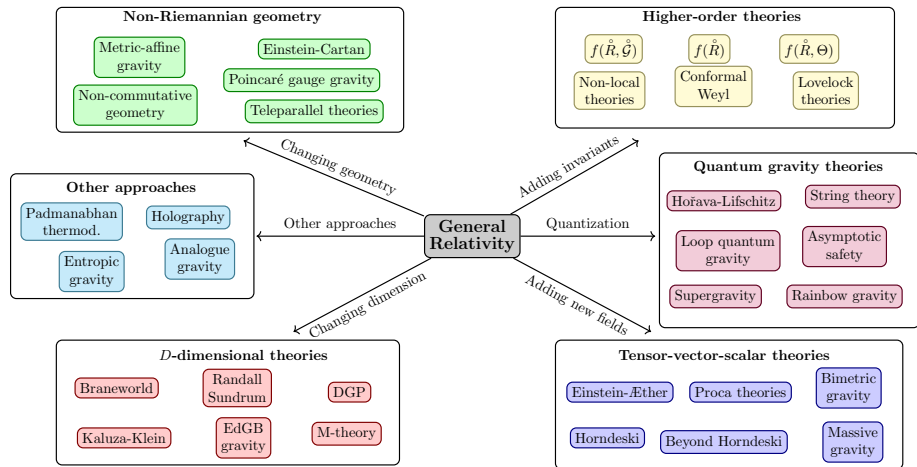


Figure: Classification of theories of gravity. (S. Bahamonde et.al., [arXiv:2106.13793 [gr-qc]].)

Gauss-Bonnet invariant

$$G = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2,$$

- Quadratic in the curvature
- $D > 4$:

$$S = \int d^D x \sqrt{-g} G$$

Leads to second-order field equations

- $D = 4$: Topological invariant \implies Pure boundary term \implies No change in the dynamics

$$G = \nabla_\mu [\Gamma^3]^\mu$$

Divergence of a non-diff-invariant vector

Theories with Gauss-Bonnet in $D = 4$

How to make the Gauss-Bonnet invariant nontrivial in $D = 4$:

- $f(G)$ theories

$$S = \int d^4x \sqrt{-g} f(G)$$

No longer second order

- scalar-Gauss-Bonnet

$$S = \int d^4x \sqrt{-g} f(\phi) G$$

belongs to Horndeski (Kobayashi, Yamaguchi, Yokoyama (2011))

⇒ Second-order field equations

$$\mathcal{L} = M_P^2 R - \frac{1}{2}(\partial\phi)^2 + M_{P\alpha}\phi G$$

- Shift-symmetric
- Horndeski ($G_5 \sim \log X$)
- Evades no-hair theorems by Hui & Nicolis (2012)

Scalar EOM around Schwarzschild

$$\square\phi = M_{P\alpha} G = \frac{12M_{P\alpha} r_s^2}{r^6}$$

Black holes

$$\phi \simeq \frac{Q_s}{r} \quad (\text{long-range})$$

with

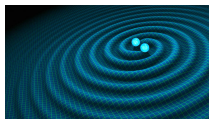
$$Q_s = \frac{M_{P\alpha}}{r_s}$$

Other sources

$$\phi \simeq \frac{M_{P\alpha} r_s^2}{r^4} \quad (\text{short-range})$$

Bounds

Observational constraint: Inspiral



Dephasing due to scalar wave emission:

$$\alpha \lesssim (1.2\text{km})^2$$

Lyu, Jiang, Yagi (2022)

Deviations from GR $< 1\%$ for black holes of $10M_{\odot}$

Causality constraint:

Absence of superluminal propagation:

$$\Lambda_{UV} < \frac{1}{\sqrt{\alpha}} \sim 1 \text{ km}^{-1}$$

Serra, Serra, Trincherini, LGT (2022)

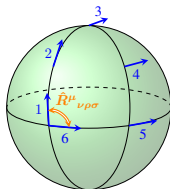
Changing geometry

Teleparallel gravity

Geometric notions

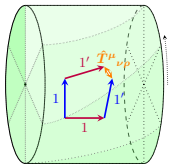
Curvature tensor $R^\alpha{}_{\beta\mu\nu}$

Rotation experienced by a vector when it is parallel transported along a closed curve



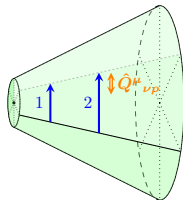
Torsion tensor $T^\alpha{}_{\beta\mu\nu}$

non-closure of the parallelogram formed when two infinitesimal vectors are parallel transported along each other.



Non-metricity tensor $Q_{\alpha\mu\nu}$

measures how much the length and angle of vectors change as we parallel transport them, so in metric spaces the length of vectors is conserved



Fundamental variables and characteristic tensors

- In the most general metric-affine setting, the fundamental variables are a **metric** $g_{\mu\nu}$ (10 comp.) as well as the coefficients $\Gamma^\rho{}_{\mu\nu}$ (64 comp.) of an **affine connection**.
- The most general connection can be written as

Connection decomposition

$$\Gamma^\lambda{}_{\mu\nu} = \underbrace{\overset{\circ}{\Gamma}^\lambda{}_{\mu\nu}}_{\text{Levi-Civita}} + \underbrace{\frac{1}{2} T^\lambda{}_{\mu\nu} - T_{(\mu}{}^\lambda{}_{\nu)}}_{\text{Torsion part: } K^\lambda{}_{\mu\nu}} + \underbrace{\frac{1}{2} Q^\lambda{}_{\mu\nu} - Q_{(\mu}{}^\lambda{}_{\nu)}}_{\text{Nonmetricity part: } L^\lambda{}_{\mu\nu}} = \overset{\circ}{\Gamma}^\lambda{}_{\mu\nu} + N^\lambda{}_{\mu\nu}$$

where we have defined the following geometrical tensors use

Curvature	$R^\mu{}_{\nu\rho\sigma} = \partial_\rho \Gamma^\mu{}_{\nu\sigma} - \partial_\sigma \Gamma^\mu{}_{\nu\rho} + \Gamma^\mu{}_{\tau\rho} \Gamma^\tau{}_{\nu\sigma} - \Gamma^\mu{}_{\tau\sigma} \Gamma^\tau{}_{\nu\rho}$
Torsion	$T^\mu{}_{\nu\rho} = \Gamma^\mu{}_{\rho\nu} - \Gamma^\mu{}_{\nu\rho}$
Nonmetricity	$Q_{\mu\nu\rho} = \nabla_\mu g_{\nu\rho} = \partial_\mu g_{\nu\rho} - \Gamma^\sigma{}_{\nu\mu} g_{\sigma\rho} - \Gamma^\sigma{}_{\rho\mu} g_{\nu\sigma}$

Curvature tensor decomposition

- The curvature becomes

$$R^\mu{}_{\nu\rho\sigma} = \overset{\circ}{R}{}^\mu{}_{\nu\rho\sigma} + \overset{\circ}{\nabla}{}_\rho N^\mu{}_{\nu\sigma} - \overset{\circ}{\nabla}{}_\sigma N^\mu{}_{\nu\rho} + N^\mu{}_{\tau\rho} N^\tau{}_{\nu\sigma} - N^\mu{}_{\tau\sigma} N^\tau{}_{\nu\rho}.$$

- Contracting the curvature tensor to obtain the Ricci scalar $R = g^{\mu\nu} R^\rho{}_{\mu\rho\nu}$ we find

Ricci scalar decomposition

$$R = \overset{\circ}{R} + \left(T + 2\overset{\circ}{\nabla}{}_\mu(\sqrt{-g}T^\rho{}_{\rho}{}^\mu) \right) + \left(Q + \overset{\circ}{\nabla}{}_\mu Q^{\mu\nu}{}_\nu - \overset{\circ}{\nabla}{}_\nu Q_\mu{}^{\mu\nu} \right) + C$$

with

$$T := T^{\rho\lambda\kappa} T_{\rho\lambda\kappa} + 2T^{\rho\lambda\kappa} T_{\kappa\rho\lambda} - 4T_{\rho}{}^{\kappa}{}_{\kappa} T^{\rho\lambda}{}_{\lambda}, \quad \text{Torsion scalar,}$$

$$Q := -\frac{1}{4} Q_{\alpha\beta\gamma} Q^{\alpha\beta\gamma} + \frac{1}{2} Q_{\alpha\beta\gamma} Q^{\beta\alpha\gamma} + \frac{1}{4} Q_\alpha Q^\alpha - \frac{1}{2} Q_\alpha \bar{Q}^\alpha, \quad \text{Nonmetricity scalar,}$$

$$C := 2(Q_{\kappa\rho\lambda} T^{\lambda\kappa\rho} + Q_\rho{}^\sigma{}_\sigma T^{\rho\kappa}{}_\kappa - Q^\sigma{}_{\sigma\rho} T^{\rho\kappa}{}_\kappa).$$

Teleparallel geometries

- **General Teleparallel geometry** ($R_{\alpha\mu\nu\beta} = 0$): In the case of vanishing curvature, the connection is flat.
- **Torsional Teleparallel geometry** ($R_{\alpha\mu\nu\beta} = 0, Q_{\alpha\mu\nu} = 0$): The metric satisfies the metric-compatibility condition but torsion is non-zero.
- **Symmetric Teleparallel geometry** ($R_{\alpha\mu\nu\beta} = 0, T^{\alpha}{}_{\mu\nu} = 0$): Both torsion tensor and curvature are zero and the gravitational interactions are only mediated through non-metricity.

Trinity of gravity

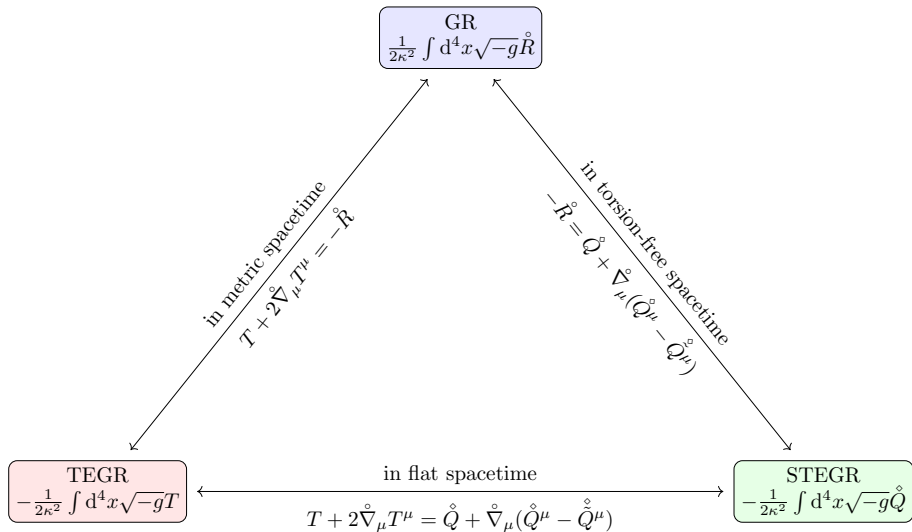


Figure: Geometrical trinity of gravity (S. Bahamonde et.al., [arXiv:2106.13793 [gr-qc]].)

Extending GR

- Modifications of GR differ depending of the geometrical formulation.

$$\overset{\circ}{R} = T + B_T = Q + B_Q$$

And then:

$$f(\overset{\circ}{R}) \neq f(T) \neq f(Q)$$

We can no longer drop the total derivative terms B_T or B_Q .

- What about Gauss-Bonnet?

In the case of [Torsional Teleparallel](#) geometry it has been shown:

Kofinas, Saridakis (2014)

$$\overset{\circ}{G} = T_G^{(T)} + B_G^{(T)}$$

Then we can construct theories like:

$$S = \int d^4x \sqrt{-g} f(T, B_T, T_G^{(T)}, B_G^{(T)})$$

or by coupling $T_G^{(T)}$ and $B_G^{(T)}$ with a scalar independently.

General teleparallel Gauss-Bonnet invariant

Riemannian Gauss-Bonnet invariant:

$$\mathring{G} = \mathring{R}_{\mu\nu\rho\sigma} \mathring{R}^{\mu\nu\rho\sigma} - 4\mathring{R}_{\mu\nu} \mathring{R}^{\mu\nu} + \mathring{R}^2$$

In differential form language (in D dimensions):

$$\mathring{G} * 1 = \frac{1}{(D-4)!} \epsilon_{a_1 \dots a_D} \mathring{R}^{a_1 a_2} \wedge \mathring{R}^{a_3 a_4} \wedge e^{a_5} \wedge \dots \wedge e^{a_D}$$

Under the teleparallel condition $R^{ab} = 0$

$$\mathring{R}^{a_1 a_2} = -\mathring{D} N^{a_1 a_2} - N^{a_1}{}_f \wedge N^{f a_2}$$

$$\begin{aligned} \mathring{G} * 1 = & \frac{1}{(D-4)!} \epsilon_{a_1 \dots a_D} \left[-d \left(N^{a_1 a_2} \wedge \mathring{R}^{a_3 a_4} \wedge e^{a_5} \wedge \dots \wedge e^{a_D} \right) \right. \\ & \left. - N^{a_1}{}_f \wedge N^{f a_2} \wedge \mathring{R}^{a_3 a_4} \wedge e^{a_5} \wedge \dots \wedge e^{a_D} \right] \end{aligned}$$

We now replace $\mathring{R}^{a_3 a_4}$ once more time in the second term,

$$\begin{aligned} \mathring{G} * 1 = & \frac{1}{(D-4)!} \epsilon_{a_1 \dots a_D} \left[-d \left(N^{a_1 a_2} \wedge \mathring{R}^{a_3 a_4} \wedge e^{a_5} \wedge \dots \wedge e^{a_D} \right) \right. \\ & \left. + \left(N^{a_1}{}_f \wedge N^{f a_2} \wedge \mathring{D} N^{a_3 a_4} + N^{a_1}{}_f \wedge N^{f a_2} \wedge N^{a_3}{}_h \wedge N^{h a_4} \right) \wedge e^{a_5} \wedge \dots \wedge e^{a_D} \right] \end{aligned}$$

Gauss-Bonnet invariant - tensorial form

$$\overset{\circ}{G} = {}^2T_G^{(T,Q)} + {}^2B_G^{(T,Q)}$$

$${}^2T_G^{(T,Q)} = \delta_{\mu_1\mu_2\mu_3\mu_4}^{\mu\nu\rho\sigma} \left[N^{\mu_1}_{\alpha\mu} N^{\alpha\mu_2}_{\nu} N^{\mu_3}_{\beta\rho} N^{\beta\mu_4}_{\sigma} - N^{\mu_1}_{\alpha\mu} N^{\alpha\mu_2}_{\nu} \overset{\circ}{\nabla}_{\sigma} N^{\mu_3\mu_4}_{\rho} \right]$$

$${}^2B_G^{(T,Q)} = -\frac{1}{2} \frac{1}{\sqrt{-g}} \partial_{\mu} \left[\sqrt{-g} \delta_{\mu_1\mu_2\mu_3\mu_4}^{\mu\nu\rho\sigma} N^{\mu_1\mu_2}_{\nu} \overset{\circ}{R}{}^{\mu_3\mu_4}_{\rho\sigma} \right]$$

where

$$\delta_{\mu_1\mu_2\mu_3\mu_4}^{\mu\nu\rho\sigma} = \frac{1}{(D-4)!} \epsilon_{\mu_1\dots\mu_D} \epsilon^{\mu\nu\rho\sigma\mu_5\dots\mu_D}$$

Gauss-Bonnet invariant - alternative split

$$\overset{\circ}{G} = {}^1T_G^{(T,Q)} + {}^1B_G^{(T,Q)}$$

$$\begin{aligned} {}^1T_G^{(T,Q)} &= \delta_{\mu_1\mu_2\mu_3\mu_4}^{\mu\nu\rho\sigma} \left[N^{\mu_1}_{\alpha\mu} N^{\alpha\mu_2}_{\nu} N^{\mu_3}_{\beta\rho} N^{\beta\mu_4}_{\sigma} - 2N^{\mu_1\mu_2}_{\mu} N^{\mu_3}_{\alpha\nu} N^{\alpha}_{\beta\rho} N^{[\beta\mu_4]_{\sigma}} \right. \\ &\quad + 2g_{\alpha\beta} N^{\mu_1\mu_2}_{\mu} N^{[\mu_3\alpha]_{\nu}} N^{\beta\mu_4}_{\gamma} N^{\gamma}_{\rho\sigma} + 2g_{\alpha\beta} N^{\mu_1\mu_2}_{\mu} N^{[\mu_3\alpha]_{\nu}} \nabla_{\sigma} N^{\beta\mu_4}_{\rho} \\ &\quad \left. + 4g_{\alpha\beta} g_{\gamma\delta} N^{\mu_1\mu_2}_{\mu} N^{[\mu_3\alpha]_{\nu}} N^{\mu_4\gamma}_{\rho} N^{(\delta\beta)_{\sigma}} \right] \end{aligned}$$

$${}^1B_G^{(T,Q)} = \frac{1}{\sqrt{-g}} \partial_{\mu} \left[\sqrt{-g} \delta_{\mu_1\mu_2\mu_3\mu_4}^{\mu\nu\rho\sigma} N^{\mu_1\mu_2}_{\nu} \left(N^{\mu_3}_{\lambda\rho} N^{\lambda\mu_4}_{\sigma} - \frac{1}{2} \overset{\circ}{R}^{\mu_3\mu_4}_{\rho\sigma} \right) \right]$$

Properties

- Known **Torsional teleparallel** expression is recovered for $N^\lambda_{\mu\nu} \rightarrow K^\lambda_{\mu\nu}$ ($Q^\lambda_{\mu\nu} \rightarrow 0$) in the first set
- In $D = 4$:

$$iT_G^{(T,Q)} = \overset{\circ}{G} - iB_G^{(T,Q)} = \text{boundary term}$$

- All pieces $iT_G^{(T,Q)}$ and $iB_G^{(T,Q)}$ contain up to second order derivatives of $g_{\mu\nu}$ and $\Gamma^\lambda_{\mu\nu}$

Symmetric teleparallel conditions

Adak et al. (2006), Beltrán Jiménez et al. (2017)

- No torsion:

$$T^{\lambda}{}_{\mu\nu} = \Gamma^{\lambda}{}_{\mu\nu} - \Gamma^{\lambda}{}_{\nu\mu} = 0 \implies \Gamma^{\lambda}{}_{\mu\nu} = (\Lambda^{-1})^{\lambda}{}_{\alpha} \partial_{\mu} \Lambda^{\alpha}{}_{\nu}$$

- No curvature:

$$R^{\mu}{}_{\nu\rho\sigma} = \partial_{\rho} \Gamma^{\mu}{}_{\nu\sigma} - \partial_{\sigma} \Gamma^{\mu}{}_{\nu\rho} + \Gamma^{\mu}{}_{\tau\rho} \Gamma^{\tau}{}_{\nu\sigma} - \Gamma^{\mu}{}_{\tau\sigma} \Gamma^{\tau}{}_{\nu\rho} = 0 \implies \Gamma^{\lambda}{}_{\mu\nu} = \frac{\partial x^{\lambda}}{\partial \xi^{\alpha}} \partial_{\mu} \partial_{\nu} \xi^{\alpha}$$

- Maximum # of d.o.f. associated to the sym. teleparallel connection = 4

- Coincident gauge:

$$\xi^{\mu} = x^{\mu} \implies \Gamma^{\lambda}{}_{\mu\nu} = 0 \implies \nabla_{\mu} = \partial_{\mu}$$

Einstein-Hilbert action becomes “Einstein action”; diff-invariance lost

$$S_{\text{STTEGR}}|_{\Gamma=0} = \frac{M_{\text{Pl}}^2}{2} \int d^4x \sqrt{-g} g^{\mu\nu} \left(\overset{\circ}{\Gamma}{}^{\alpha}{}_{\beta\mu} \overset{\circ}{\Gamma}{}^{\beta}{}_{\nu\alpha} - \overset{\circ}{\Gamma}{}^{\alpha}{}_{\beta\alpha} \overset{\circ}{\Gamma}{}^{\beta}{}_{\mu\nu} \right)$$

Symmetric TG Gauss-Bonnet scalars in FRLW

- Consider flat FLRW background
- Ensure the scalar, metric and connection respect the same symmetries: Killing vector fields Z_ζ where $\zeta = \{1..m\}$. Solve

$$\mathcal{L}_{Z_\zeta} \phi = 0 \quad (\mathcal{L}_{Z_\zeta} g)_{\mu\nu} = 0 \quad (\mathcal{L}_{Z_\zeta} \Gamma)^\lambda{}_{\mu\nu} = 0$$

$$\implies ds^2 = -N(t)^2 dt^2 + a(t)^2 (dr^2 + r^2 d\vartheta^2 + r^2 \sin^2 \vartheta d\varphi^2)$$

For the connection part, it is convenient to decompose the metric as

$$g_{\mu\nu} = -n_\mu n_\nu + h_{\mu\nu}, \quad n_\mu = (-N, 0, 0, 0)$$

$$\implies Q_{\rho\mu\nu} = 2F_1(t) n_\rho n_\mu n_\nu + 2F_2(t) n_\rho h_{\mu\nu} + 2F_3(t) h_{\rho(\mu} n_{\nu)}$$

- Impose the teleparallel condition (flat connection) \implies 3 distinct branches

Branch 1: $F_1 = K, F_2 = -H, F_3 = 0$

Branch 2: $F_1 = 2H + \frac{1}{KN} \frac{dK}{dt}, F_2 = -H, F_3 = K$

Branch 3: $F_1 = -K - \frac{1}{K} \frac{dK}{dt}, F_2 = K - H, F_3 = K$

Symmetric TG Gauss-Bonnet scalars in FRLW

$$\overset{\circ}{G} = {}^1T_G^{(Q)} + {}^1B_G^{(Q)} = {}^2T_G^{(Q)} + {}^2B_G^{(Q)} = 24H^2 \left(\frac{\dot{H}}{N} + H^2 \right).$$

First set

$${}^1T_G^{(Q)} = \begin{cases} 24H^2 \left(\frac{\dot{H}}{N} + H^2 \right), & \text{First branch} \\ \frac{6(2H+K)^2\dot{H}}{N} + \frac{12H(H+K)\dot{K}}{N} + 6H^2(6HK+4H^2+3K^2), & \text{Second branch} \\ \frac{6(K-2H)^2\dot{H}}{N} + \frac{12H(K-H)\dot{K}}{N} + 6H^2(-6HK+4H^2+3K^2), & \text{Third branch} \end{cases}$$

$${}^1B_G^{(Q)} = \begin{cases} 0, & \text{First branch} \\ -\frac{6K(4H+K)\dot{H}}{N} - \frac{12H(H+K)\dot{K}}{N} - 18H^2K(2H+K), & \text{Second branch} \\ \frac{6K(4H-K)\dot{H}}{N} + \frac{12H(H-K)\dot{K}}{N} + 18H^2K(2H-K), & \text{Third branch} \end{cases}$$

Second set

$${}^2T_G^{(Q)} = \begin{cases} -12H^2 \left(\frac{\dot{H}}{N} + H^2 \right), & \text{First branch} \\ -\frac{12H(H+K)\dot{H}}{N} - \frac{6H^2\dot{K}}{N} - 6H^3(2H+3K), & \text{Second branch} \\ \frac{12H(K-H)\dot{H}}{N} + \frac{6H^2\dot{K}}{N} + 6H^3(3K-2H), & \text{Third branch} \end{cases}$$

$${}^2B_G^{(Q)} = \begin{cases} 36H^2 \left(\frac{\dot{H}}{N} + H^2 \right), & \text{First branch} \\ \frac{12H(3H+K)\dot{H}}{N} + \frac{6H^2\dot{K}}{N} + 18H^3(2H+K), & \text{Second branch} \\ \frac{6H(6H-2K)\dot{H}}{N} - \frac{6H^2\dot{K}}{N} + 18H^3(2H-K). & \text{Third branch} \end{cases}$$

Symmetric TG scalar-Gauss-Bonnet

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left[Q - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + \alpha_1 \mathcal{G}_1(\phi) T_G^{(Q)} + \alpha_2 \mathcal{G}_2(\phi) B_G^{(Q)} \right].$$

- Riemannian scalar-Gauss-Bonnet is contained as a special case:

$$\alpha_1 \mathcal{G}_1(\phi) = \alpha_2 \mathcal{G}_2(\phi)$$

- Second order field equations in all the fields $(\phi, g_{\mu\nu}, \Gamma^\alpha{}_{\mu\nu})$ for *both* sets
- For the **second set only**, the theory can be written with up to quadratic contractions of $Q_{\lambda\mu\nu} \implies$ It belongs to the **Symmetric Teleparallel Horndeski** class

Bahamonde, Trenkler, LGT, Yamaguchi (2022)

Summary

- Theories including the Gauss-Bonnet invariant have interesting phenomenology
- Alternative geometrical formulations of gravity provide new ways to modify GR
- We extended the Gauss-Bonnet invariant to General Teleparallel constructions, in particular the Symmetric Teleparallel case.
- We found that the splitting between a bulk and a boundary term is not unique, and in $D = 4$ both pieces are boundary terms.
- In a flat FLRW background in Symmetric Teleparallel formulation 3 branches of solutions exist. The corresponding ST-Gauss-Bonnet invariants include the d.o.f. of the connection $K(t)$ nontrivially.
- Example theory: Scalar-ST-Gauss-Bonnet theory has second-order field equations.
- Future directions include:
 - Cosmological evolution
 - Scalarized BHs in scalar-ST-Gauss-Bonnet
 - Generalization to Metric-Affine Gravity
 - ...

Thank you for your attention!